

Solutions to PS 2 Physics 201

1.

$$\mathbf{E} = \mathbf{i} \int \frac{k_e dq}{r^2} \quad (1)$$

$$= \mathbf{i} \int_{-L}^L \frac{k_e \lambda_0 x dx}{L(x_0 - x)^2} \quad (2)$$

$$= \mathbf{i} \frac{k_e \lambda_0}{L} \int_{-L}^L \frac{(x - x_0) dx}{(x - x_0)^2} + \frac{x_0 dx}{(x - x_0)^2} \quad (3)$$

$$= \mathbf{i} \frac{k_e \lambda_0}{L} \left[\ln \left(\frac{x_0 - L}{x_0 + L} \right) + \frac{x_0}{x_0 - L} - \frac{x_0}{x_0 + L} \right] \quad (4)$$

$$= \mathbf{i} \frac{k_e \lambda_0}{L} \left[\ln \left(\frac{x_0 - L}{x_0 + L} \right) + \frac{2x_0 L}{x_0^2 - L^2} \right] \quad (5)$$

To find the field for $x_0 \rightarrow \infty$, we first want to rewrite this result in terms of the small parameter $\frac{L}{x_0}$. Doing so yields

$$\mathbf{E} = \mathbf{i} \frac{k_e \lambda_0}{L} \left[\ln \left(\frac{1 - \frac{L}{x_0}}{1 + \frac{L}{x_0}} \right) + \frac{2 \frac{L}{x_0}}{1 - \left(\frac{L}{x_0} \right)^2} \right] \quad (6)$$

Next, we perform a Taylor expansion in terms of $\frac{L}{x_0}$ about the point $\frac{L}{x_0} = 0$. For the logarithm term, we find

$$\ln \left(\frac{1 - \frac{L}{x_0}}{1 + \frac{L}{x_0}} \right) \approx 0 - \frac{2L}{x_0} - \frac{2L^3}{3x_0^3} \quad (7)$$

and for the second term

$$\frac{2x_0 L}{x_0^2 - L^2} = 2 \frac{L}{x_0} \left[\frac{1}{1 - \left(\frac{L}{x_0} \right)^2} \right] \quad (8)$$

$$= 2 \frac{L}{x_0} \left[1 + \left(\frac{L}{x_0} \right)^3 \right] \quad (9)$$

$$= \frac{2L}{x_0} + \frac{2L^3}{x_0^3} \quad (10)$$

Putting this all together, we arrive at a final approximation for \mathbf{E} given by

$$\mathbf{E} = \mathbf{i} \frac{k_e \lambda_0}{L} \left(-\frac{2L}{x_0} - \frac{2L^3}{3x_0^3} + \frac{2L}{x_0} + \frac{2L^3}{x_0^3} \right) \quad (11)$$

$$= \mathbf{i} \frac{k_e \lambda_0}{L} \frac{4L^3}{3x_0^3} \quad (12)$$

$$= \mathbf{i} \frac{\lambda_0 L^2}{3\pi \epsilon_0 x_0^3} \quad (13)$$

as desired. Comparing this to the expression for a dipole field aligned with the axis of a dipole, we find

$$\mathbf{i} \frac{2p}{4\pi \epsilon_0 x_0^3} = \mathbf{i} \frac{\lambda_0 L^2}{3\pi \epsilon_0 x_0^3} \quad (14)$$

$$p = \frac{2\lambda_0 L^2}{3} \quad (15)$$

2.

$$\boldsymbol{\tau} = \mathbf{p} \times \mathbf{E} \quad (16)$$

$$= -pE \sin \theta \mathbf{k} \quad (17)$$

$$= \mathbf{k}(10^{-29})(0.5) \sin \frac{\pi}{6} N \cdot m \quad (18)$$

$$= \mathbf{k}2.5 \times 10^{-30} N \cdot m \quad (19)$$

To find the work done, we use

$$W_{done} = -W_e = \Delta U \quad (20)$$

$$= U(\pi) - U\left(\frac{\pi}{6}\right) \quad (21)$$

$$= -pE \cos \pi + pE \cos \frac{\pi}{6} \quad (22)$$

$$= pE \left(1 + \frac{\sqrt{3}}{2} \right) \quad (23)$$

$$= 5 \left(1 + \frac{\sqrt{3}}{2} \right) \times 10^{-30} N \cdot m \quad (24)$$

Finally, for the frequency of small oscillations, we use Newton's second law

$$\tau = \frac{dL}{dt} \quad (25)$$

$$-pE \sin \theta = I\ddot{\theta} \quad (26)$$

$$(27)$$

expanding $\sin \theta$ for small θ , we find

$$-\frac{pE}{I}\theta = \ddot{\theta} \quad (28)$$

$$\omega^2 = \frac{pE}{I} \quad (29)$$

$$(30)$$

Plugging in the numbers, we have

$$I = m \left(\frac{d}{2}\right)^2 + m \left(\frac{d}{2}\right)^2 \quad (31)$$

$$= \frac{md^2}{2} \quad (32)$$

$$= 5 \times 10^{-48} \text{ kg} \cdot \text{m}^2 \quad (33)$$

and thus

$$\omega = \sqrt{\frac{pE}{I}} \quad (34)$$

$$= \sqrt{\frac{5 \times 10^{-30} \text{ rad}}{5 \times 10^{-48} \text{ s}}} \quad (35)$$

$$= 10^9 \frac{\text{rad}}{\text{s}} \quad (36)$$

3. By spherical symmetry, we know automatically that the electric field everywhere will be purely in the radial direction. Using this fact, we can apply Gauss' law by finding the flux through a sphere of radius r centered about the origin,

$$\int \mathbf{E} \cdot d\mathbf{A} = 4\pi r^2 E_r = \frac{Q_{\text{enclosed}}}{\epsilon_0} \quad (37)$$

For $r < a$, we have that

$$Q_{\text{enclosed}} = -\frac{\frac{4}{3}\pi r^3 Q}{\frac{4}{3}\pi a^3} \quad (38)$$

$$= -Q \frac{r^3}{a^3} \quad (39)$$

and therefore

$$\mathbf{E}(r < a) = -\mathbf{e}_r \frac{k_e Q r}{a^3} \quad (40)$$

For $a \leq r < b$, we have that $Q_{enclosed} = -Q$, and so

$$\mathbf{E}(a \leq r < b) = -\mathbf{e}_r \frac{k_e Q}{r^2} \quad (41)$$

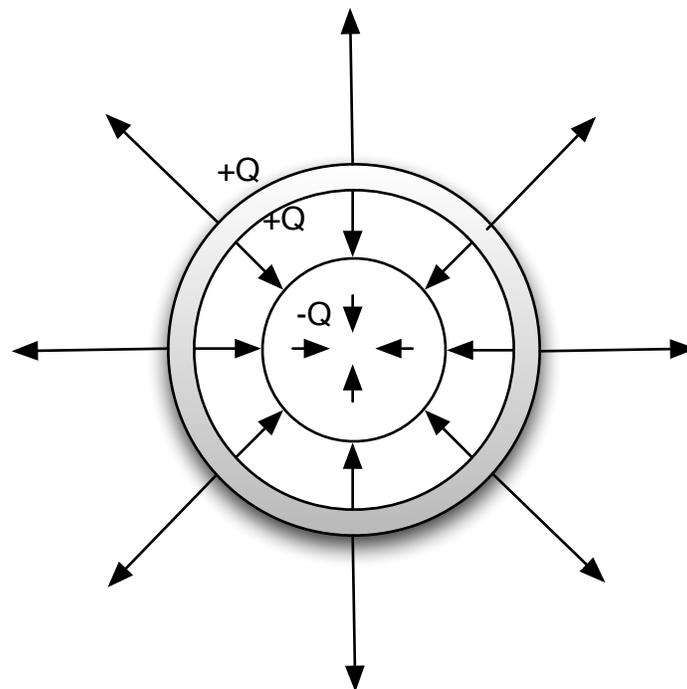
For $b \leq r < c$, the field must be zero since this describes the interior of a conductor, meaning that a charge of $+Q$ must reside on the interior surface of the conducting shell. Therefore

$$\mathbf{E}(b \leq r < c) = 0 \quad (42)$$

Finally, for $r \geq c$, it must be that $Q_{enclosed} = +Q$, and therefore

$$\mathbf{E}(r \geq c) = \mathbf{e}_r \frac{k_e Q}{r^2} \quad (43)$$

Below is a sketch showing where the charges reside, and some field lines.



4. Exploiting the cylindrical symmetry of the problem tells us that the field directed radially (i.e. in the \mathbf{e}_r direction) away from the axis of the cylinders, and that

$$2\pi r L E_r = \frac{Q_{enclosed}}{\epsilon_0} \quad (44)$$

where L is the length of our cylindrical Gaussian surface.

Thus, since the cylinders are hollow, we know that there is no charge enclosed for $r < a$, and thus

$$\mathbf{E}(r < a) = 0 \quad (45)$$

For $a \leq r < b$, we have that

$$Q_{\text{enclosed}} = \lambda L \quad (46)$$

and therefore

$$\mathbf{E}(a \leq r < b) = \mathbf{e}_r \frac{2k_e \lambda}{r} \quad (47)$$

Lastly, for $r \geq b$, we have that $Q_{\text{enclosed}} = 0$, so

$$\mathbf{E}(r \geq b) = 0 \quad (48)$$

To find the surface charge density σ on the inner cylinder, we note that we can express the total charge on a length L of the cylinder as either

$$Q = 2\pi a L \sigma \quad (49)$$

or as

$$Q = \lambda L \quad (50)$$

Equating these two expressions, we find that

$$\sigma = \frac{\lambda}{2\pi a} \quad (51)$$

We can substitute this result into our expression for the field between the cylinder to find

$$\mathbf{E}(a < r < b) = \mathbf{e}_r \frac{4\pi k_e a \sigma}{r} \quad (52)$$

$$= \mathbf{e}_r \frac{a \sigma}{\epsilon_0 r} \quad (53)$$

For $b - a \ll a$, we have that between the cylinders $r - a \ll a$, and thus

$$\mathbf{E}(a < r < b) = \mathbf{e}_r \frac{a \sigma}{\epsilon_0 r} \quad (54)$$

$$= \mathbf{e}_r \frac{a \sigma}{\epsilon_0 (a + r - a)} \quad (55)$$

$$\approx \mathbf{e}_r \frac{\sigma}{\epsilon_0} \quad (56)$$

This is equal in magnitude to the field of a parallel plate capacitor of the same charge density. Furthermore, on a very small scale \mathbf{e}_r does not vary significantly with the polar angle, and thus may be approximated as a cartesian unit vector. Thus, this setup locally approximates a parallel plate capacitor

5. By Gauss' law,

$$\Phi_e = \frac{1C}{\epsilon_0} \quad (57)$$

Thus, by the symmetry of the cube, we must have that the flux through one of the faces is given by

$$\Phi_e = \frac{1C}{6\epsilon_0} = 1.88 \times 10^{10} \frac{N \cdot m^2}{C} \quad (58)$$

6. For $r < R$ we have that

$$Q_{enclosed} = \int_0^r \int_0^\pi \int_0^{2\pi} \rho(r)r^2 \sin \theta d\phi d\theta d\rho \quad (59)$$

$$= 4\pi A \int_0^r r^4 dr \quad (60)$$

$$= \frac{4\pi A}{5} r^5 \quad (61)$$

Evaluating this same expression at $r = R$ gives that the total charge Q is

$$Q = \frac{4\pi A}{5} R^5 \quad (62)$$

and thus for $r < R$

$$Q_{enclosed} = Q \frac{r^5}{R^5} \quad (63)$$

Thus, Gauss' law tells us that

$$\mathbf{E}(r < R) = \mathbf{e}_r \frac{k_e Q r^3}{R^5} \quad (64)$$

For $r > R$, we have that $Q_{enclosed} = Q$, and therefore

$$\mathbf{E}(r \geq R) = \mathbf{e}_r \frac{k_e Q}{r^2} \quad (65)$$

7. By the Pythagorean theorem, the radius of each disc as a function of z is given by

$$r(z) = \sqrt{R^2 - z^2} \quad (66)$$

and thus the area A of each disc is

$$A(z) = \pi (R^2 - z^2) \quad (67)$$

The volume of the sphere is then given by integrating over all discs contained in the sphere, i.e. from $z = -R$ to $z = R$. This yields

$$V = \pi \int_{-R}^R (R^2 - z^2) dz \quad (68)$$

$$= \pi \left(2R^3 - \frac{2R^3}{3} \right) \quad (69)$$

$$= \frac{4}{3} \pi R^3 \quad (70)$$

as desired

8. Knowing that Gauss' law follows from Coulomb's law, we can define an analogue of Gauss's law for the gravitational field \mathbf{G} . Examining the form of both Newton's and Coulomb's law, we have

$$\mathbf{G} = -\frac{GM}{r^2} \mathbf{e}_r \quad (71)$$

and

$$\mathbf{E} = \frac{q}{4\pi\epsilon_0 r^2} \mathbf{e}_r \quad (72)$$

where the fields \mathbf{E} and \mathbf{G} are the forces on a unit charge (unit mass) due to charge q (mass M). By comparison, we see that M plays the same roll as q , and likewise $-G$ plays the same roll as $\frac{1}{4\pi\epsilon_0}$. From this we arrive at Gauss's law for gravitation,

$$\int \mathbf{G} \cdot d\mathbf{A} = -4\pi GM_{enclosed} \quad (73)$$

9. Using Gauss' law, we have that for $r = .5m$, $Q_{enclosed} = 1\mu C$, and thus

$$\mathbf{E} = \hat{r} \frac{k_e(1\mu C)}{(.5m)^2} = 3.6 \times 10^4 \frac{N}{C} \mathbf{e}_r \quad (74)$$

Next, for $r = 2m$, $Q_{enclosed} = -1\mu C$, and thus

$$\mathbf{E} = -\hat{r} \frac{k_e(1\mu C)}{(2m)^2} = -2.2 \times 10^3 \frac{N}{C} \mathbf{e}_r \quad (75)$$

10. We can examine this situation as a solid sphere of uniform charge density ρ and radius R superimposed with a solid sphere of uniform charge density $-\rho$ with radius $\frac{R}{2}$. Let \mathbf{r}_1 denote the vector from the center of the larger sphere to a point within the smaller sphere, and let \mathbf{r}_2 denote the vector from the center of the smaller sphere to that same point.

First, we use Gauss' law to find the field \mathbf{E}_+ due to the larger sphere. At a distance r_1 , we have that the charge enclosed is given by

$$Q_{\text{enclosed}} = \frac{4}{3}\pi r_1^3 \rho \quad (76)$$

and thus the field \mathbf{E}_+ is given by

$$\mathbf{E}_+ = \frac{\rho r_1}{3\epsilon_0} \mathbf{e}_{r_1} = \frac{\rho}{3\epsilon_0} \mathbf{r}_1 \quad (77)$$

Similarly, for the field E_- due to the smaller, negatively charged sphere, we find

$$\mathbf{E}_- = -\frac{\rho r_2}{3\epsilon_0} \mathbf{e}_{r_2} = -\frac{\rho}{3\epsilon_0} \mathbf{r}_2 \quad (78)$$

Summing together these two contributions to find the total field in the cavity, we get

$$\mathbf{E} = \mathbf{E}_+ + \mathbf{E}_- = \frac{\rho}{3\epsilon_0} (\mathbf{r}_1 - \mathbf{r}_2) \quad (79)$$

But from the figure, we can see that

$$\mathbf{r}_1 - \mathbf{r}_2 = \frac{R}{2} \hat{\mathbf{x}} \quad (80)$$

Thus,

$$\mathbf{E} = \frac{\rho R}{6\epsilon_0} \hat{\mathbf{x}} \quad (81)$$

which describes a uniform field in the $\hat{\mathbf{x}}$ direction