

PS# 5 due Thursday:

NOTE: orbital period =
8 hrs on prob 1
test after break
(test prep sheet upcoming
review session 9:15pm night
~~test~~ before)

POST-NEWTONIAN EFFECTS of G.R.

1) precession of perihelion
(Mercury's orbit)

periastron
perigalacticon

2) deflection of light
eclipse observations
→ gravitational lenses

3) gravitational redshift

$$z = \frac{\Delta \lambda}{\lambda} = \frac{\Delta P_p}{P_p}$$

↑
"redshift"

$$= \frac{1}{\sqrt{1 - R_s/r}} - 1$$

Schw. radius of
some mass

distance of
the light
source from
mass

how much redshift
do you observe
from infinite distance

$$R_s/r \rightarrow 0$$

if observer
is not at
infinity

$$z_{obs} = z_{source \rightarrow \infty} - z_{observer \rightarrow \infty}$$

prob
set #2

has some redsh.ft



D

has less redsh.ft

change in redsh.ft

4) gravitational waves

effect: orbital
period
decreases

NOT observed in
solar system

LAB FOR G.R.
"Binary Pulsar"

pulsars: discovered in late
1960s

observed: pulses of radio
waves

pulsation happens ~~over~~
with period

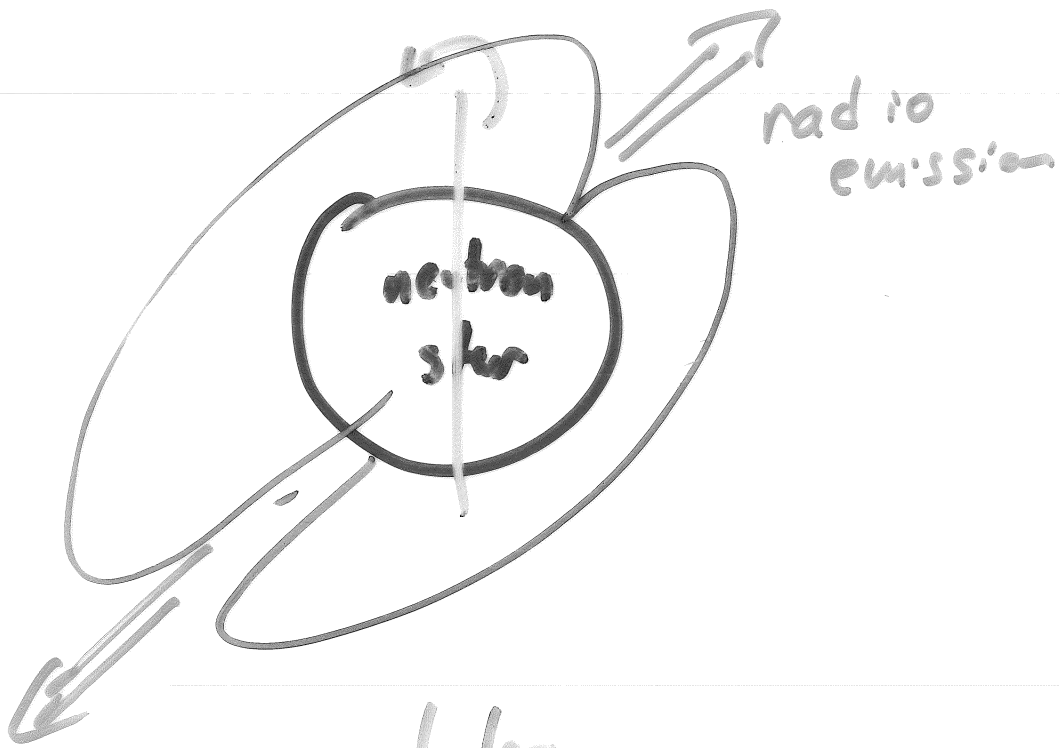
between milliseconds
to tens of seconds

rotating magnetized
neutron stars

FABLE: Jocelyn Bell and
discovery of pulsars

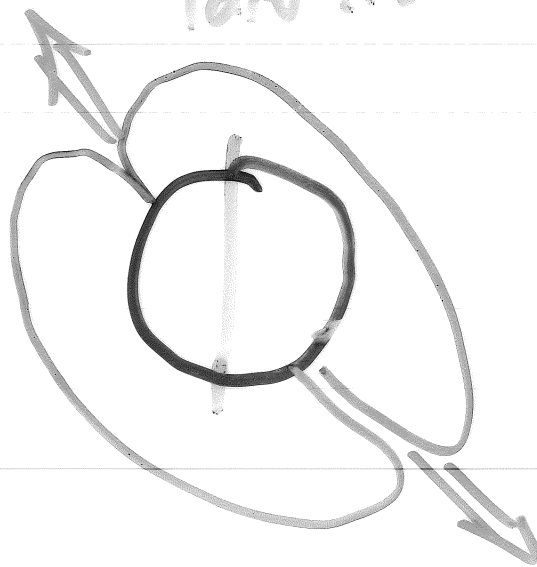
MORAL: grad students } never get
women } the credit

: Thinking and doing are
not the same



ON

later...

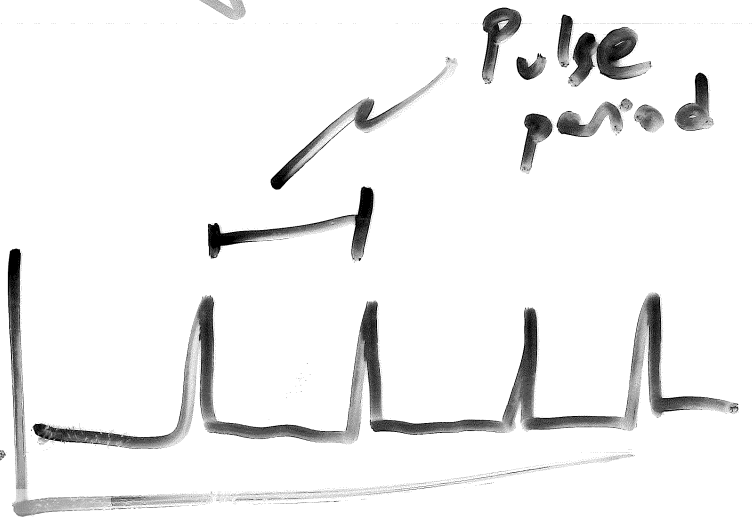


OFF

Pulse period

⇒

radio emission



time

$$\frac{\Delta P_p}{P_p} = \sqrt{\frac{1 + v_R/c}{1 - v_R/c}} - 1$$

$$\approx \frac{v_R}{c} \rightarrow \text{Newton approx}$$

Discovery of binary pulsar

PERIODICAL ≈ 8 hrs

measure pulse period to

high accuracy

elliptical orbit

precession of periastron

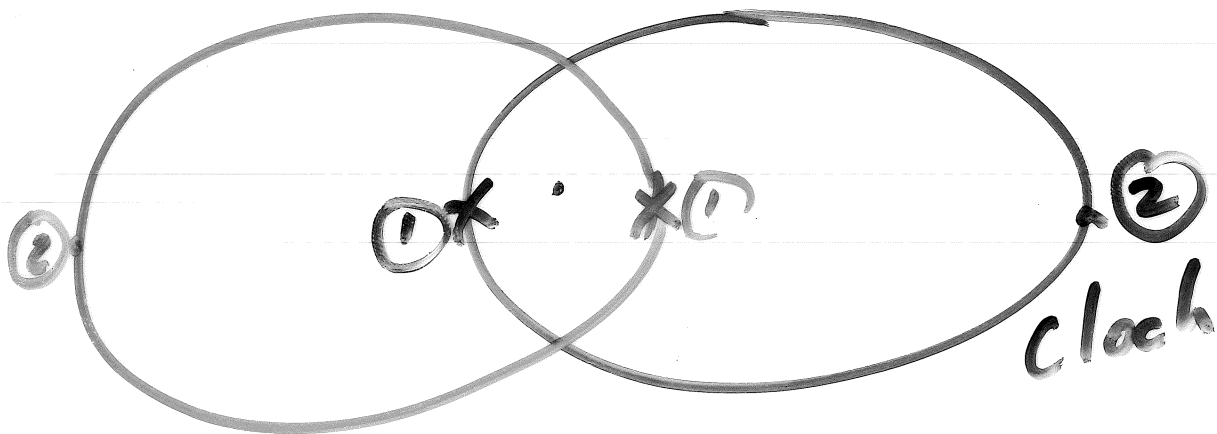
4° / year

Open Yale courses

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(Mercury: $43''$ / century)

also: grav. redsh.ft
seeing effect of the
companion object
(also neutron star)



Lower
object

at position ① objects
close redshift caused by
other ~~star~~ object is large

at position ② distance
is greater \rightarrow redshift is

less

ALSO: orbital period
decreases due to
grav. waves

only Newtonian measurements

→ determine M_{TOT}
if you know inclination

measure precession of periastron
also depends on $M_1, M_2,$
and i

SOLVE for inclination
know M_{TOT} but not
each individual mass

grav. redshift:
depends on M_1, M_2, c

how: solve for
 M_1 (also M_2)

know M_1, M_2 , inclination

PREDICT what the
period change should
be

measure this effect

→ prediction correct
to measurement errors

CLEAR that G.R. is

"correct" in post-Newtonian
approximation

BUT imagine a
theory which is like
Newton (1st term)
like G.R. (2nd term
"post-Newtonian")

but different at very
strong grav. fields

so it would be good
to test STRONG FIELD
effect
 $R_s/r \rightarrow$ close to 1