

PS #4 pls hand in
ON CHAIRS

$$(d\tau)^2 = \frac{(dr)^2}{(1 - \frac{R_s}{r})^{\frac{R_s}{r}}} + r^2(d\Omega)^2 - c^2(1 - \frac{R_s}{r})(dt)^2$$

r is a coordinate
 $R_s = 2GM/c^2$
 $\rightarrow \Omega, \tau$

object of M , R_0 \rightarrow radius of object

black hole: $R_0 < R_s$

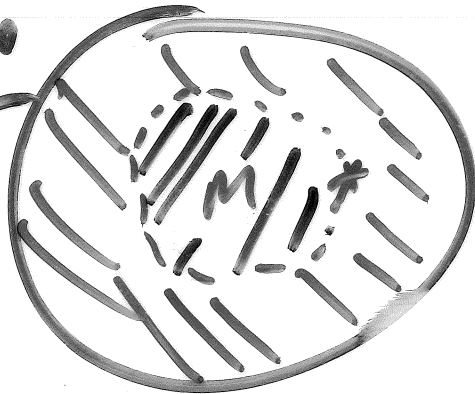
\hookrightarrow some r where
 $r = R_s$

if $R_0 > R_s$

Then no $r = R_s$

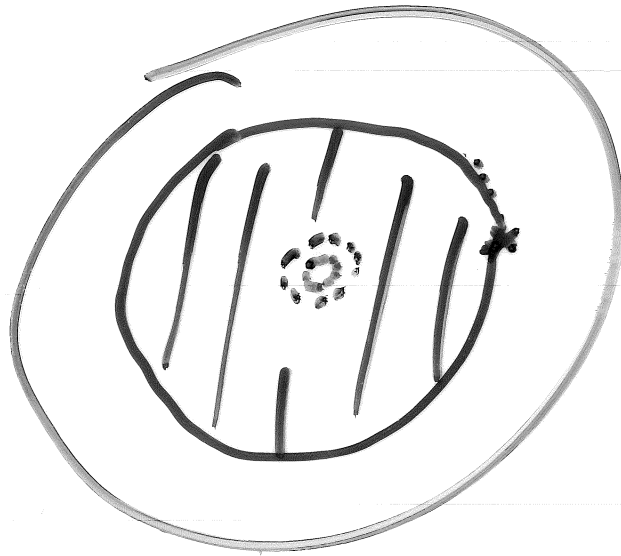
The relevant mass is
mass INSIDE r

object
with M, R_0



So for $r < R_0$ \rightarrow radius of object
then relevant $M < M_0$ \rightarrow total mass of object

if $R_0 > R_s$
then there is no
 $r < R_s$



$$R_0 > R_S$$

$$M$$

What is R_S
 that goes into the
 metric for person x

$$R_0 < R_S$$



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$r = R_S$ **ARGH!**

POST-NEWTONIAN GRAVITATION EFFECTS

P-N effect #1:
precession of perihelion
(e.g. Mercury)

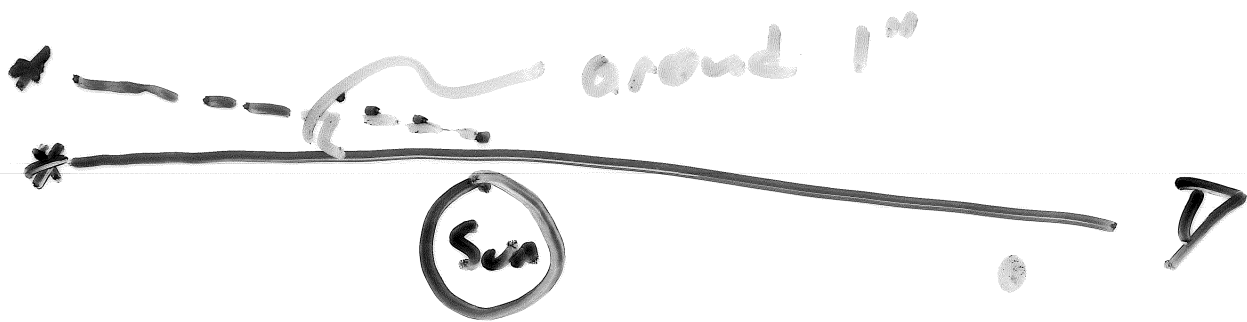
P-N effect #2:
deflection of light



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dist. galaxy, quasar

* in solar system



look at stars behind ~~sun~~
appears to be in a
different position than
when observed at night.

DO THIS BY OBSERVING
DURING ECLIPSE

1917 Einstein publishes G.R.

1919 eclipse in Brazil

Eddington mounts

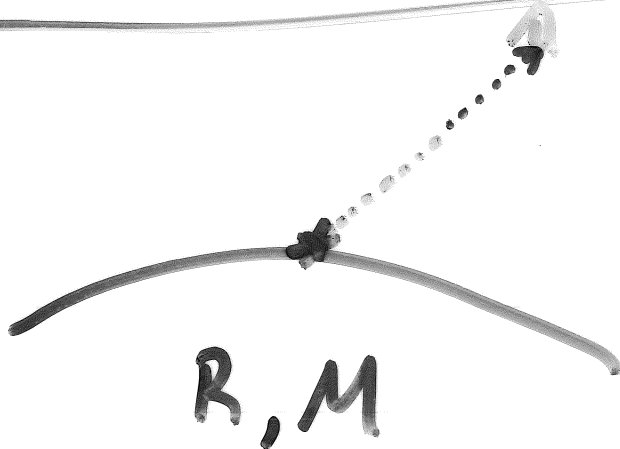
expedition to Brazil/
WORKS

FABLE: 1919 eclipse expedition

MORAL: Science is intentional
universal ~~post~~-quest

: Science works as
advised
maybe?

post-N effect #3
gravitational redshift

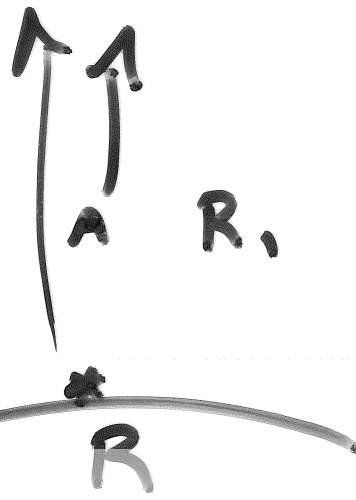
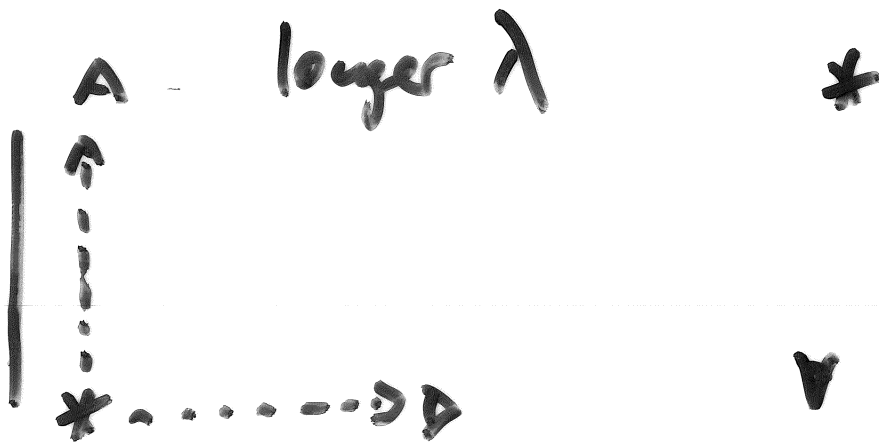


$$\frac{\Delta\lambda}{\lambda} = \frac{1}{\sqrt{1 - R_s/R}} - 1$$

distance to MASS

E of photon
 gets less \nearrow
 $= \frac{hc}{\lambda}$ speed of light
 \rightarrow longer
 \rightarrow Planck's constant

TESTED IN LABS



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$$\frac{\Delta \lambda}{\lambda_0} = \frac{\Delta R_1}{R_0} - \frac{\Delta \lambda_R}{\lambda_0}$$

P-N effect #4:

gravitational waves

as mass moves back & forth
(in orbit)

→ ripples in space-time
propagate outwards at
c

The energy in "gravitational
waves"

comes from orbit

orbit gets gradually
smaller

objects spiral in

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Can see this in binary
stars with short orbit
"BINARY PULSAR"